## Present status of $\theta_{13}$



## Solar neutrinos prefer $\theta_{13} \neq 0$



## Reactor and Accelerator experiments don't see any signal

Reactors: CHOOZ (1998)



## Reactor and Accelerator experiments don't see any signal



Backgrounds:  $49.1 \pm 7(\text{stat}) \pm 2.7(\text{syst})$ 



## How to build the signal/exclusion plot (II)

- The minimum of the  $\chi^2$  distribution is the best fit
- The region at a given confidence level (CL) is defined by the contour at a given  $\Delta\chi^2$  from the minimum.
- The CL is computed from the probability distribution of a  $\chi^2$  at two degrees of freedom  $(\sin^2 2\theta, \Delta m^2)$

**Question:** Why  $\Delta \chi^2$  and not  $\chi^2$ ? **Hint:** Why two degrees of freedom?

A more formal approach in G.Feldman and R.Cousins, Phys.Rev.D57:3873-3889,1998



## $\theta_{13}$ at reactors



- $P_{\bar{e}\bar{e}}$  and  $\theta_{13}$ , directly connected, no controbution by  $\delta_{\rm CP}$  e sign $(\Delta m_{23}^2)$ .
- No way to measure CP violation and mass hierarchy
- Complementary to the accelerators.
- Disappearance experiments: systematic errors dominate.



## **Reactor Experiments**

# CHOOZ result $R = 1.01 \pm 2.8\%(\text{stat}) \pm 2.7\%(\text{syst})$ .

Goal:

• Improve by a factor 5, at least, the statistical error (25 times more neutrinos)

- Bigger detectors (CHOOZ was 5 ton fiducial)
- More stable in time (CHOOZ took data for 8761.7 h, then stopped because liquid scintillator degraded thanks to the gadolinium.)

• Improve systematic errors by a factor 5 at least:

- Add a close detector
- Design a detector optimized to a better background reduction and a better control of systematics.

Nuclear reactors are a very intense source of  $\overline{\nu}_e$  from  $\beta$  decays of the fission fragments.

Every fission reaction emits about 200 MeV of energy and 6  $\overline{\nu}_e$ .  $\Downarrow$ Flux  $\sim 2 \cdot 10^{20} \ \overline{\nu}_e \ s^{-1} \ \text{GWatt}^{-1}$ , isotropic,  $\langle E(\overline{\nu}_e) \rangle \simeq 0.5 \ \text{MeV}$ .

Oscillation experiments look for  $\overline{\nu}_e$  disappearance at different baselines:

•  $L = O(1 \text{km}) \Rightarrow$  atmospheric regime: Double Chooz, RENO, Daya Bay.

•  $L = O(200 \text{km}) \Rightarrow$  solar regime: Kamland

## Neutrino flux

See Bemporad, Gratta, Vogel Rev.Mod.Phys.74:297,2002

Detect absolute number of neutrino interaction and distortions of their spectrum

prompt positron signal, energy range.  $\overline{
u_e} p \rightarrow e^+ n$   $n + p \xrightarrow[\tau \simeq 186 \ \mu s]} d + \gamma (2.2 \ MeV)$ delayed correlated photon.

To determine neutrino flux:

- Measure of the reactor thermal power
- 2 Determination of the neutrino spectrum
- Definition of the experimental observable: positron momentum spectrum.



## Thermal power of the reactor

The leading reaction is  $^{235}U$  fission:

 $^{235}U + n \rightarrow X_1 + X_2 + 2n$ 

The lightest fragment have on average  $A \simeq$  94, the heavier:  $A \simeq$  140. Stable nuclei with A = 94, 140 are  ${}_{40}Zr^{94}$  e  ${}_{58}Ce^{140}$ .  ${}^{235}U$  has 98 protons and 142 neutrons  $\Rightarrow$  to reach the stability, on average it needs 6 neutron  $\beta$  decays  $\Rightarrow$  6  $\overline{\nu}_e$ .



The interaction process  $\overline{\nu}_e + p \rightarrow n + e^+$  has a threshold of  $\sim 1.8 \ MeV \Rightarrow$  only  $\sim 25\%$  of neutrinos can be detected.

All the neutrinos from low Q-value processes, as nuclear fuel stored in the reactors and radioactivity induced in the nuclear plant structures, don't produce detectable neutrinos.

The fuel composition of the reactor core changes with the time, it's under monitor (reactor power depends from its composition).



### From fission rate to the $\overline{\nu}_e$ spectrum

The  $\overline{\nu}_e$ spectrum of three of the four principal fission nuclei: (<sup>235</sup>U, <sup>239</sup>Pu, <sup>241</sup>Pu), has been derived by measuring the electron spectrum. The fourth: <sup>238</sup>U, has been computed from nuclear models, as well all the processes in the decay chain. Systematic error:  $\sim 1\%$ .

From  $\overline{\nu}_e$  to positrons  $\overline{\nu}_e + p \rightarrow n + e^+$  cross section:  $\sigma_{tot}^{(0)} = \sigma_0 (f^2 + 3g^2) E_e^{(0)} p_e^{(0)}$  $= 0.0952 \left( \frac{E_e^{(0)} p_e^{(0)}}{1 \text{ MeV}^2} \right) \times 10^{-42} \text{ cm}(1)$ 

 $E_e^{(0)} = E_{\nu} - (M_n - M_p)$ : positron energy (neglecting neutron recoil, marginal effect)  $p_e^{(0)}$  momentum,

f = 1, g = 1.26 vector and axial coupling constants

$$\sigma_0 = \frac{G_F^2 \cos^2 \theta_C}{\pi} (1 + \Delta_{inner}^R) , \qquad (2)$$

radiative corrections:  $\Delta^R_{inner} \simeq 0.024$ .



Solid lines: predictions  $atO(1/M_n)$ , dashed O(1).

## Data/prediction agreement

Experiment Bugey 3 (years 80', now considered a non oscillation experiment): expected and measured  $\overline{\nu}_e$ spectrum.

Curve b) is the most updated prediction.



## Systematic

errors summary (from hep-ph/0107277) Origin and magnitude of systematic errors in PALO VERDE and CHOOZ. Note that the two experiments offer different breakdowns of their systematics. For simplicity we do not show the systematics for the PALO VERDE ON-OFF analysis. The PALO VERDE results are from the analysis of the full data set (Boehm *et al.* 2001).

Systematic	Chooz (	(%)	P.V.	(%)
$\sigma(\overline{\nu}_{\rm e} + {\rm p}  ightarrow {\rm n} + {\rm e}^+)$	1.9		-	
Number of p in target	0.8		-	
W <sub>th</sub>	0.7		-	
Energy abs. per fission	0.6		-	
Total rate prediction	2	2.3		2.1
e <sup>+</sup> trigger eff.	-		2.0	
n trigger eff.	-		2.1	
$\overline{\nu}_e$ selection cuts	-		2.1	
$(1-\epsilon_1)B_{\rm pn}$ estimate	-		3.3	
Total $\overline{\nu}_e$ efficiency	1	5		4.9
Total	2	2.7		5.3



### Two main categories:

- Accidental backgrounds from the random superposition of a "positron-like" and "neutron-like" signals. Directly estimated from the measured rates of the two processes.
- Backgrounds from neutrons induced by cosmic rays. They can be measured only if the reactor is off (impossible to pay to have a reactor shutdown).



Chooz counting rate as function of the reactors power.



Inner detector Liquid scintillator doped with Gadolinium (0.1%) in a acrylic vessel. Gadolinium increases neutron capture cross section, reducing capture time from  $\sim$  170  $\mu s$  to  $\sim$  27  $\mu s \Rightarrow$  with a reduction of not-correlated noise. Furthermore it increases the energy of the  $\gamma s$  produced by the neutron capture, from  $\sim$  2 MeV to  $\sim$  8 MeV.



Inner detector Liquid scintillator doped with Gadolinium (0.1%) in a acrylic vessel. Gadolinium increases neutron capture cross section, reducing capture time from  $\sim$  170  $\mu s$  to  $\sim$  27  $\mu s \Rightarrow$  with a reduction of not-correlated noise. Furthermore it increases the energy of the  $\gamma s$  produced by the neutron capture, from  $\sim$  2 MeV to  $\sim$  8 MeV.

Gamma catcher: Not doped liquid scintillator to capture the  $\gamma s$  emitted by the neutron capture. It allows a better definition of the fiducial volume.



Inner detector Liquid scintillator doped with Gadolinium (0.1%) in a acrylic vessel. Gadolinium increases neutron capture cross section, reducing capture time from  $\sim$  170  $\mu$ s to  $\sim$  27  $\mu$ s  $\Rightarrow$  with a reduction of not-correlated noise. Furthermore it increases the energy of the  $\gamma s$  produced by the neutron capture, from  $\sim$  2 MeV to  $\sim$  8 MeV.

Gamma catcher: Not doped liquid scintillator to capture the  $\gamma$ s emitted by the neutron capture. It allows a better definition of the fiducial volume.

**Phototubes shield**: Not scintillating oil to separate the active volume from the phototubes, the most important source of radioactivity in the detector.



Inner detector Liquid scintillator doped with Gadolinium (0.1%) in a acrylic vessel. Gadolinium increases neutron capture cross section, reducing capture time from  $\sim$  170  $\mu$ s to  $\sim$  27  $\mu$ s  $\Rightarrow$  with a reduction of not-correlated noise. Furthermore it increases the energy of the  $\gamma s$  produced by the neutron capture, from  $\sim$  2 MeV to  $\sim$  8 MeV.

Gamma catcher: Not doped liquid scintillator to capture the  $\gamma$ s emitted by the neutron capture. It allows a better definition of the fiducial volume.

**Phototubes shield**: Not scintillating oil to separate the active volume from the phototubes, the most important source of radioactivity in the detector.

**Inner veto**: To shield against Comptons induced by external radioactivity and by crossing muons. Equipped with phototubes.



Inner detector Liquid scintillator doped with Gadolinium (0.1%) in a acrylic vessel. Gadolinium increases neutron capture cross section, reducing capture time from  $\sim$  170  $\mu s$  to  $\sim$  27  $\mu s$   $\Rightarrow$  with a reduction of not-correlated noise. Furthermore it increases the energy of the  $\gamma s$  produced by the neutron capture, from  $\sim$  2 MeV to  $\sim$  8 MeV.

Gamma catcher: Not doped liquid scintillator to capture the  $\gamma$ s emitted by the neutron capture. It allows a better definition of the fiducial volume.

**Phototubes shield**: Not scintillating oil to separate the active volume from the phototubes, the most important source of radioactivity in the detector.

**Inner veto**: To shield against Comptons induced by external radioactivity and by crossing muons. Equipped with phototubes.

**Outer veto**: To veto crossing muons. It is required a minimum 100 m.w.e depth to keep dead times below 25%. Some ions produced by  $\mu$ s: <sup>8</sup>He and <sup>9</sup>Li, withd decaying times of 119 ms and 174 ms cannot be vetoed anyway.



## The three players

Setup	$P_{\mathrm{Th}}$ [GW]	<i>L</i> [m]	$m_{ m Det}$ [t]	Events/year	Backgrounds/day
Daya Bay	17.4	1700	80	$10 \cdot 10^4$	0.4
Double Chooz	8.6	1050	8.3	$1.5\cdot 10^4$	3.6
RENO	16.4	1400	15.4	$3 \cdot 10^4$	2.6





# **Double Chooz**

### Talk by J. Dawson



### 2 cores - 1 site - 8.5 GW<sub>th</sub>

### 1 near position, 1 far

- target: 2 x 8.3 t
   Civil engineering
- 1 near lab ~ Depth 40 m, Ø 6 m
- 1 available lab

### Statistics (including ɛ)

- far: ~ 40 evts/day
- near: ~ 460 evts/day

#### Systematics

- reactor : ~ 0.2%
- detector : ~ 0.5%

### Backgrounds

- $\sigma_{b2b}$  at far site: ~ 1%
- $\sigma_{h2h}$  at near site: ~ 0.5%

### Planning

- 1. Far detector only
- Sensitivity (1.5 ans) ~ 0.06
- 2. Far + Near sites
  - available from 2010
  - Sensitivity (3 years) ~ 0.025

# Daya Bay



## RENO



Mauro Mezzetto (INFN Padova)

## Reactors systematic business

Error Description	CHOOZ	Double Chooz			Daya Bay		
·					No R&D R&D		
	Absolute	Absolute	Relative	Absolute	Relative	Relative	
Reactor							
Production cross section	1.90 %	1.90 %		1.90 %			
Core powers	0.70 %	2.00 %		2.00 %			
Energy per fission	0.60 %	0.50 %		0.50 %			
Solid angle/Bary. displct.			0.07 %		0.08 %	0.08 %	
Detector							
Detection cross section	0.30 %	0.10 %		0.10 %			
Target mass	0.30 %	0.20 %	0.20 %	0.20 %	0.20 %	0.02 %	
Fiducial volume	0.20 %						
Target free H fraction	0.80 %	0.50 %		?	0.20 %	0.10 %	
Dead time (electronics)	0.25 %						
Analysis (paticle id.)							
$e^+$ escape (D)	0.10 %						
$e^+$ capture (Ć)							
$e^+$ identification cut (E)	0.80 %	0.10 %	0.10 %				
n escape (D)	0.10 %						
n capture (% Gd) (C)	0.85 %	0.30 %	0.30 %	0.10 %	0.10 %	0.10 %	
n identification cut (É)	0.40 %	0.20 %	0.20 %	0.20 %	0.20 %	0.10 %	
$\overline{\nu}_{e}$ time cut (T)	0.40 %	0.10 %	0.10 %	0.10 %	0.10 %	0.03 %	
$\overline{\nu}_e$ distance cut (D)	0.30 %						
unicity (n multiplicity)	0.50 %				0.05 %	0.05 %	
Total	2.72 %	2.88 %	0.44 %	2.82 %	0.39 %	0.20 %	

G. Mention, T. Lasserre and D. Motta, arXiv:0704.0498 [hep-ex].

## Reactors vs Accelerators: 2018



## What to compare

**Sensitivity**: The highest value of a parameter (say  $\theta_{13}$ ) that can be excluded at a given CL in absense of a signal. True value = 0; fit value  $\neq 0$ .

**Discovery potential**: The smallest value of a parameter that can be provide a signal that can't be fitted with a null value at a given CL.

True value  $\neq$  0; fit value = 0.

**Probability of discovery**: The probability that a given true value of  $\theta_{13}$  produces a not null result at a given CL

## Discovery potential at 3 $\sigma$ : 2018



## Sensitivity time evolution



## Discovery Potential: time evolution



# Probability of excluding $\theta_{13} = 0$ (3 $\sigma$ )

